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Search for Heavy Neutrinos in $K^+ \to \mu^+ \nu_\mu$ Decays

The NA62 Collaboration*

The NA62 experiment recorded a large sample of $K^+ \to \mu^+ \nu_{\mu}$ decays in 2007. A peak search has been performed in the reconstructed missing mass spectrum. In the absence of a signal, limits in the range 2×10^{-6} to 10^{-5} have been set on the squared mixing matrix element $|U_{\mu4}|^2$ between muon and heavy neutrino states, for heavy neutrino masses in the range 300–375 MeV/ c^2 . The result extends the range of masses for which upper limits have been set on the value of $|U_{\mu4}|^2$ in previous production search experiments.

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1 Introduction

The fact that neutrinos oscillate implies that they have non-zero masses. While in the Standard Model (SM) neutrinos are massless by construction, the SM can be extended in various ways to accommodate neutrino masses [1]. In a large class of models, the see-saw mechanism is used to explain the lightness of the SM neutrinos by introducing additional heavy neutrino mass states which mix with the SM flavour states [2]. One example of models including heavy neutrinos is the neutrino minimal Standard Model (ν MSM), in which three right-handed neutrinos are added to the SM with one of them being at the GeV scale [3, 4]. For heavy neutrinos with masses below the kaon mass, limits on their mixing matrix elements can be placed by searching for peaks in the missing mass spectrum of K^{\pm} decays [5]. In the following, two-body kaon decays to a muon and a SM neutrino are denoted $K^+ \to \mu^+\nu_{\mu}$, while those with a muon and a heavy neutrino are denoted $K^+ \to \mu^+\nu_h$; the notation $K^+ \to \mu^+\nu_h$ indicates either case. Limits on $|U_{\mu4}|^2$ in the extended neutrino mixing matrix using the process $K^+ \to \mu^+\nu_h$ come from experiments with stopped kaons, and are of the order of 10^{-8} up to 300 MeV/ c^2 [6] and 10^{-6} up to 330 MeV/ c^2 [7].

The ratio of the K^+ decay width to heavy neutrino to the decay width to SM muon neutrinos is related to $|U_{\mu4}|^2$ [5]:

$$\frac{\mathcal{B}(K^+ \to \mu^+ \nu_h)}{\mathcal{B}(K^+ \to \mu^+ \nu_\mu)} = |U_{\mu 4}|^2 f(m_h) \,, \tag{1}$$

where m_h is the mass of the heavy neutrino, and $f(m_h)$ accounts for the phase space factor and the helicity suppression, and varies in the range 1.5–4.0 for m_h in the region 300–375 MeV/ c^2 considered in the present analysis.

Under the assumption that heavy neutrinos decay only to SM particles, the lifetime of a heavy neutrino is determined by the mixing matrix elements and by its mass [8]. For heavy neutrino masses in the range 300–375 MeV/ c^2 , the dominant decay modes are $\nu_h \to \pi^0 \nu_{e,\mu,\tau}$ and $\nu_h \to \pi^+ \ell^-$, where $\ell = e, \mu$. Assuming $|U_{\ell 4}|^2 < 10^{-4}$ with $\ell = e, \mu, \tau$, the mean free path of heavy neutrinos at NA62 for any mass in the range considered is greater than 10 km, and therefore their decays can be neglected, since the probability of decaying in the detector or decay volume is below 1%.

2 Beam, detector and data samples

The beam line and detector of the earlier NA48/2 experiment were reused by the NA62 experiment during 2007 data taking; they are described in detail in [9, 10]. Primary protons of 400 GeV/c, extracted from the CERN SPS, impinged on a 40 cm long, 0.2 cm diameter beryllium target. Secondary beams of positively and negatively charged hadrons were produced, momentum-selected, similarly focused and transported to the detector. These beams could be run simultaneously or separately. The central beam momentum of 74 GeV/c was selected by the first two magnets in a four-dipole achromat and by momentum-defining slits incorporated into a 3.2 m thick copper/iron proton beam dump, which also provided the possibility of blocking either of the two beams. The beams had a momentum spread of $\pm 1.4 \text{ GeV}/c$ (rms). For about 1.8×10^{12} primary protons incident on the target per SPS spill of 4.8 s duration, the secondary beam fluxes at the entrance to the decay volume were, respectively, 1.7×10^7 and 0.8×10^7 positively and negatively charged particles per spill.

The fraction of kaons in each beam was about 6%. The beam kaons decayed in a fiducial volume contained in a 114 m long cylindrical evacuated tank. The K^+ and K^- beams were deflected at the entrance of the fiducial volume by angles in the range $\pm (0.23 - 0.30)$ mrad with respect to the longitudinal z axis to compensate for the opposite ∓ 3.58 mrad deflections by the

spectrometer magnet. All these deflections were regularly reversed during the data taking. The hadron beams were accompanied by a flux of stray (halo) muons produced by kaon and pion decays upstream of the decay volume. Two magnetized iron toroids with apertures centred on the beam line were installed upstream of the decay volume to deflect positive halo muons and suppress the associated backgrounds.

Charged particle momenta were measured by a magnetic spectrometer, housed in a tank filled with helium at approximately atmospheric pressure, placed downstream of the decay volume. The spectrometer consisted of four drift chambers (DCHs), each comprising four views of double planes of staggered sense wires, and a dipole magnet located between the second and the third DCHs. The magnet provided a horizontal transverse momentum kick of 265 MeV/c to charged particles, and the spectrometer had a momentum resolution of $\sigma_p/p = 0.48\% \oplus 0.009\% \cdot p$, where the momentum p is expressed in GeV/c. A hodoscope (HOD) consisting of two planes of plastic scintillator strips producing fast trigger signals was placed after the spectrometer. A liquid krypton (LKr) electromagnetic calorimeter of thickness $127 \text{ cm} (27X_0)$ was located further downstream. Its 13248 readout cells had a transverse size of 2×2 cm² each with no longitudinal segmentation. The energy resolution was $\sigma_E/E = 3.2\%/\sqrt{E} \oplus 9\%/E \oplus 0.42\%$, and the spatial resolution for the transverse coordinates x and y of an isolated electromagnetic shower was $\sigma_x = \sigma_y = 0.42 \text{ cm} / \sqrt{E} \oplus 0.06 \text{ cm}$, where E is expressed in GeV. A muon detector (MUV) was located further downstream. The MUV was composed of three planes of plastic scintillator strips (aligned horizontally in the first and last planes, and vertically in the middle plane) read out by photomultipliers at both ends. Each strip was 2.7 m long and 1 cm thick. The widths of the strips were 25 cm in the first two planes, and 45 cm in the third plane. The MUV was preceded by a hadronic calorimeter (6.7 nuclear interaction lengths) not used for the present measurement. Each MUV plane was preceded by an additional 0.8 m thick iron absorber.

General data taking conditions are described in [11]. The main trigger condition for selecting the sample of $K^+ \to \mu^+ N$ decays required the coincidence in time and space of signals in the two HOD planes (HOD signal), and loose lower and upper limits on the DCH hit multiplicity (1-track signal), downscaled by a factor of 150. Data taking periods with simultaneous beams were collected with a lead bar installed between the two HOD planes for muon identification studies. For data collected with the lead bar in place, the vetoing power for backgrounds with photons is reduced, so these data are excluded from the present analysis. Since the muon halo background is smaller in the K^+ sample, this analysis is based on data with the K^+ beam only (43% of the integrated kaon flux, as used in [12]), while data taken with only the K^- beam are used to study the background from halo muons.

3 Analysis strategy

In the decay $K^+ \to \mu^+ N$ the neutrino mass can be reconstructed as $m_h^2 = m_{\text{miss}}^2 = (p_K - p_\mu)^2$, where p_K and p_μ are the four-momenta of the kaon and the muon respectively. The kaon momentum is not measured on an event-by-event basis, and p_K is obtained, assuming the kaon mass, from the average three-momentum measured with $K^+ \to \pi^+ \pi^+ \pi^-$ decays approximately every 500 SPS spills. The muon four-momentum p_μ is determined as that of a reconstructed charged track, assumed to be a muon.

Simulated samples of $K^+ \to \mu^+ \nu_h$ decays, with mass m_h between 240 and 380 MeV/ c^2 at 1 MeV/ c^2 intervals, have been generated to determine the signal acceptance. The kaon decay modes that contribute to the background have been simulated to determine the expected spectrum of the reconstructed m_{miss} . The contribution to the background from muon halo is evaluated using a control data sample, defined as the sample recorded with the K^- beam only. The integrated μ^+ halo flux in the control sample is three times smaller than that in the K^+ sample.

A comparison between the expected and observed $m_{\rm miss}$ spectra is used to set limits on the observed number of $K^+ \to \mu^+ \nu_h$ decays for each assumed ν_h mass. These are translated into limits on the branching ratio $\mathcal{B}(K^+ \to \mu^+ \nu_h)$ and the mixing parameter $|U_{\mu4}|^2$. The search for heavy neutrinos is restricted to the mass range 300–375 MeV/ c^2 (called the signal region in the following). Lower neutrino masses are accessible, but a strong limit on $|U_{\mu4}|^2$ of the order of 10^{-8} from a production experiment exists below 300 MeV/ c^2 [7]. The reconstructed missing mass range 245–298 MeV/ c^2 , where the ν_h presence is excluded by this limit, is used as a control region to measure the trigger efficiency for the background events.

4 Event selection

Charged particle trajectories and momenta are reconstructed from hits and drift times in the spectrometer using a detailed magnetic field map. The reconstructed $K^+ \to \pi^+ \pi^+ \pi^-$ invariant mass is used for fine calibration of the spectrometer momentum scale and DCH alignment throughout the data taking. Clusters of energy deposition in the LKr calorimeter are found by locating maxima in space and time in the digitized pulses from individual cells. Reconstructed energies are corrected for energy outside the cluster boundaries, energy lost in isolated inactive cells (0.8% of the total number), sharing of energy between clusters, and non-linearity for clusters with energy below 11 GeV.

The selection requires exactly one positively charged track with the following characteristics: within the DCH, LKr calorimeter and MUV geometrical acceptance; momentum p between 10 and 65 GeV/c; within 20 ns of the trigger time recorded by the HOD; distance of closest approach (CDA) between the track and the beam axis, as monitored with $K^+ \rightarrow \pi^+\pi^-\pi^-$ decays, smaller than 3 cm; track extrapolation associated in time and space with MUV signals from the first two planes.

Selected events are required to be free of clusters of energy deposition in the LKr calorimeter except for any of the following configurations: the cluster energy is lower than 2 GeV; the cluster time is more than 12 ns away from the track time; the cluster is consistent with bremsstrahlung from the track before deflection by the spectrometer magnet (within 6 cm of the straight-line extrapolated upstream track); the cluster position is within 40 cm of the extrapolated downstream track.

5 Background contributions

The background receives contributions from muon halo, evaluated with the control data sample, and from kaon decays, evaluated with simulation.

Muon halo background

A data driven approach is used in modelling the muon halo contribution, and in designing a selection that minimizes this background while preserving signal acceptance. The distribution of halo background events is estimated using the control sample (see Section 3). The majority of reconstructed μ^+ in the control sample comes from muon halo with two sources of contamination: 1) K^+ in specific momentum bands pass through the beam absorbers (with a probability of up to 5×10^{-4} depending on momentum) and decay into $K^+ \to \mu^+ \nu_{\mu}$; a simulation shows that the reconstructed m_{miss} calculated assuming the nominal kaon momentum is lower than 280 MeV/ c^2 , and therefore this component does not enter the signal region; 2) the contribution from misidentified positively charged pions from $K^- \to \pi^- \pi^- \pi^+$ decays enters the signal region, and is simulated and subtracted. To study the halo, the event selection described in Section 4 is used with a relaxed CDA condition (CDA < 8 cm). The distribution of the events in the control sample passing this selection is shown in Figure 1 in the variables z_{vertex} , track momentum p and θ , where θ is the angle between the K^+ beam axis and the measured muon direction. To minimize the halo contribution, additional selection criteria are applied in a five-dimensional space ($z_{\text{vertex}}, \theta, p, \text{CDA}, \phi$), where ϕ is the track azimuthal angle in the transverse plane. The contours in Figure 1 show example projections of these five-dimensional criteria; the events outside the contours are rejected. The signal acceptance reduction due to the multi-dimensional criteria with respect to the selection described in Section 4 is in the range 40–45% depending on m_h .

The estimated number of halo background events in the final sample is obtained from the number of events observed in the control sample, normalized to the K^+ data in the range $m_{\rm miss}^2 > 0.05 {\rm ~GeV}^2/c^4$ and $3 < {\rm CDA} < 8 {\rm ~cm}$.

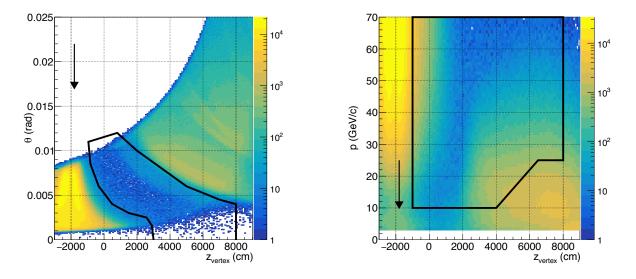


Figure 1: Distribution of halo events in the $(z_{\text{vertex}}, \theta)$ plane (left) and (z_{vertex}, p) plane (right). The contours show the projections of the five-dimensional selection criteria. The events outside the contours are rejected. The arrow indicates the start of the fiducial volume. See Section 5 for details.

Kaon decay background

The total number of kaon decays in the fiducial region, N_K , is used to scale the simulated distributions of the expected backgrounds. It is measured with a sample of $K^+ \to \mu^+ \nu_{\mu}$ decays using the selection described in [11] after adding the kinematic criteria; the number of events in the missing mass squared distribution within $|m_{\text{miss}}^2| < 0.015 \text{ GeV}^2/c^4$ is evaluated after subtracting a sub-percent contribution from beam halo. The squared missing mass distribution is shown in Figure 2. The number of $K^+ \to \mu^+ \nu_{\mu}$ decays after background subtraction is 9.45 × 10⁶ and the corresponding acceptance is 24.88%. The resulting number of kaon decays in the fiducial volume in the analysed dataset is $N_K = (5.977 \pm 0.015) \times 10^7$.

the fiducial volume in the analysed dataset is $N_K = (5.977 \pm 0.015) \times 10^7$. The decay $K^+ \to \mu^+ \nu_{\mu}$ forms a peak at zero $m_{\rm miss}^2$ with a width determined by the width of the kaon momentum spectrum, the beam divergence and the spectrometer resolution; the peak is well outside the 300–375 MeV/ c^2 signal region. The contribution from $K^+ \to \mu^+ \nu_{\mu} \gamma$ decay appears as a high-mass tail in the $m_{\rm miss}^2$ distribution and is taken into account by the simulation. The dominant background from kaon decays in the signal region comes from $K^+ \to \pi^0 \mu^+ \nu_{\mu}$ decays with an undetected π^0 due to the non-hermetic geometrical acceptance. The hadronic decay $K^+ \to \pi^+ \pi^0$ is only reconstructed as signal if the π^0 is undetected and the π^+ is misidentified as a muon or decays into a muon.

The backgrounds due to kaon decays to three pions are naturally suppressed because they involve either three tracks $(K^+ \to \pi^+ \pi^+ \pi^-)$ or photons $(K^+ \to \pi^+ \pi^0 \pi^0)$. The events which pass the selection typically appear at the upper end of the m_{miss}^2 spectrum. Decays with positrons in the final state $(K^+ \to e^+ \nu_e, K^+ \to \pi^0 e^+ \nu_e)$ are rejected with particle identification.

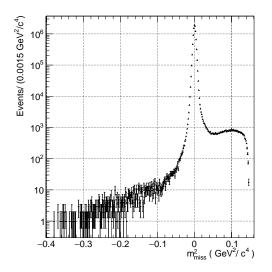


Figure 2: Reconstructed squared missing mass distribution for data passing the final event selection.

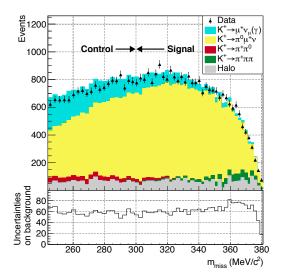


Figure 3: Missing mass distributions for data, showing statistical uncertainties, and for the estimated background contributions, in both signal and control regions. The lower plot shows the total uncertainty on the background estimate.

6 Systematic uncertainties on the background estimate

The uncertainty on kaon decay background receives contributions from the uncertainty on the number of kaon decays in the fiducial volume, N_K , and the individual kaon decay branching ratios. The contribution from the finite size of the simulated samples is negligible. The estimate of the systematic uncertainty associated with N_K is obtained by varying the cut on $|m_{\text{miss}}^2|$ by $\pm \sigma_{m^2}$ resulting in a variation of 0.2% (where $\sigma_{m^2} = 3.1 \times 10^{-3} \text{ GeV}^2/c^4$ is the resolution on the $K^+ \to \mu^+ \nu_{\mu}$ signal peak). The contribution from $\mathcal{B}(K^+ \to \mu^+ \nu_{\mu})$ results in a variation of 0.15%. The overall systematic uncertainty on kaon decay background varies from 0.6% to 1.0% of the total expected background as a function of m_{miss}^2 , and it is dominated in the signal region by the uncertainty on $\mathcal{B}(K^+ \to \pi^0 \mu^+ \nu)$.

To estimate the uncertainty on the $K^+ \to \mu^+ \nu_{\mu}$ background due to non-Gaussian tails in the DCH resolution, a sample of $K^+ \to \pi^+ \pi^0$ ($\pi^0 \to \gamma \gamma$) decays, selected with LKr calorimeter without the spectrometer information, is used. The expected π^+ three-momentum is computed from the photon cluster energies and positions using the average kaon momentum and π^0 mass constraint. Data and simulated $K^+ \to \pi^+ \pi^0$ are then compared in terms of the agreement between the measured and expected missing mass, π^+ momentum and direction spectra. From this comparison it is inferred that the uncertainty on the background estimate in the $K^+ \to \mu^+ \nu_h$ signal region does not exceed 6% of the total expected background; this uncertainty affects mostly the low ν_h mass region.

The systematic uncertainty attributed to the halo background arises from the limited size of the control sample (halo statistical contribution), and from the assumption that the halo distribution in the control sample accurately reproduces that of the K^+ data (halo model contribution). The halo statistical contribution is 2–4% of the total expected background in the range 300–360 MeV/ c^2 and rises to 16% in the range 360–375 MeV/ c^2 . The control sample is divided into sub-samples according to selection variables and each sub-sample is normalised to the K^+ data. The halo model contribution is evaluated by comparing the normalizations obtained with the different sub-samples with that obtained for the entire sample. This contribution is 1–3% of the total expected background in the range 300–360 MeV/ c^2 and rises to 8% in the range 360–375 MeV/ c^2 . The uncertainty due to the subtraction of $K^- \to \pi^- \pi^- \pi^+$ events is negligible.

A $K^+ \to \mu^+ \nu_{\mu}$ sample is used to measure the MUV muon identification efficiency as a function of track momentum. This efficiency varies between 96% and 98% over the momentum range between 10 and 65 GeV/c. The simulation is tuned to reproduce this efficiency to 1% precision, and therefore a systematic uncertainty of 1% is assigned to the total expected background.

The HOD trigger inefficiency is $(1.4\pm0.1)\%$ as discussed in [11]; since the inefficiency depends mainly on the number of tracks which is the same for signal, $K^+ \rightarrow \mu^+ \nu_{\mu}$ decays and main backgrounds, it cancels out to a good approximation. The 1-track trigger inefficiency for $K^+ \rightarrow$ $\mu^+ \nu_{\mu}$ decays was measured with respect to the HOD trigger to be much smaller that the HOD inefficiency [11] and can be neglected. Conversions of undetected photons from $K^+ \rightarrow \pi^0 \mu^+ \nu$ decays cause a 1-track inefficiency due to events with high multiplicity of hits in the DCH chambers. The 1-track trigger efficiency for the background could be evaluated directly in the signal region, or by extrapolating the measurement performed in the control region to the signal region. However, the possible presence of $K^+ \rightarrow \mu^+ \nu_h$ decays in the signal region would increase the apparent efficiency, thereby affecting the signal sensitivity. Therefore the 1-track trigger efficiency for the background is evaluated in the control $m_{\rm miss}$ region 245–298 MeV/ c^2 , since strong limits on the heavy neutrino production in this region already exist. In this control region the 1-track efficiency is (89.8±0.6)% and was shown not to depend on the missing mass. The uncertainty on the trigger efficiency for the background.

7 Upper limits on heavy neutrino production

The event selection described in Section 4 with the addition of the five-dimensional criteria described in Section 5 constitutes the final selection. Figure 3 shows the $m_{\rm miss}$ distribution of events passing the final selection and the estimated background spectrum. The halo contribution varies as a function of $m_{\rm miss}$ between 5% and 20% of the background and carries the largest relative systematic uncertainty.

For each neutrino mass m_h under consideration in the signal region 300–375 MeV/ c^2 , a window of $\pm \sigma_h$ in the missing mass spectrum is defined centred on m_h , where σ_h is the resolution parametrized as $\sigma_h = 12 \text{ MeV}/c^2 - 0.03 \cdot m_h$. For each window, the width is rounded to the nearest multiple of $10^{-4} \text{ GeV}^2/c^4$ in m_{miss}^2 . The signal acceptance, evaluated for a range of heavy neutrino masses with simulation, is about 0.20 up to 360 MeV/ c^2 and drops to zero for larger masses. The statistical analysis is performed by applying the Rolke-Lopez method [13] to find the 90% confidence intervals on the number of reconstructed $K^+ \to \mu^+ \nu_h$ events for the case of a Poisson process in the presence of Gaussian backgrounds. Inputs to the computation in each mass window are the number of data events observed, and the estimate of the total number of background events with its uncertainty. The squared uncertainties on the numbers of expected events in each mass hypothesis are shown in Figure 4, where the various contributions

can be seen.

No signal is observed, the maximum local significance being 2.67 standard deviations at 357 MeV/ c^2 . The upper limits (UL) at 90% CL on the numbers of reconstructed $K^+ \to \mu^+ \nu_h$ events is indicated as n_{UL} . The expected upper limits are calculated assuming that the number of events observed is equal to the number of events expected, i.e. the number of background events. These upper limits are converted to upper limits on the branching ratio $\mathcal{B}(K^+ \to \mu^+ \nu_h)$ as shown in Figure 5, using the relation $n_{UL} = \mathcal{B}_{UL}(K^+ \to \mu^+ \nu_h)A(m_h)N_K$, where $A(m_h)$ is the signal acceptance, \mathcal{B}_{UL} is the upper limit on the branching ratio, and N_K is given in Section 5. The branching ratio is related to the neutrino mixing-matrix element squared $|U_{\mu4}|^2$ by equation (1). The obtained upper limits on $|U_{\mu4}|^2$ are shown in Figure 6, together with the limits from a previous peak search experiment [6].

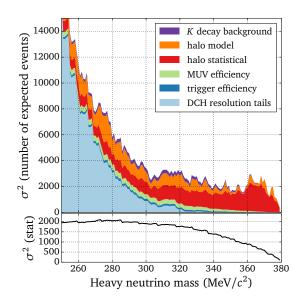


Figure 4: Upper plot: squared uncertainties on the numbers of expected background events at each heavy neutrino mass. Lower plot: squared statistical uncertainty for data.

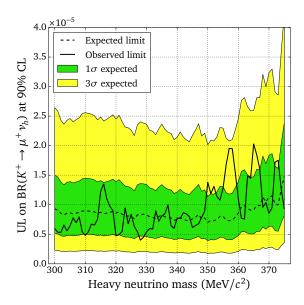


Figure 5: Expected and observed upper limits (at 90% CL) on the branching ratio in 10^{-5} units of the $K^+ \rightarrow \mu^+ \nu_h$ decay at each assumed ν_h mass.

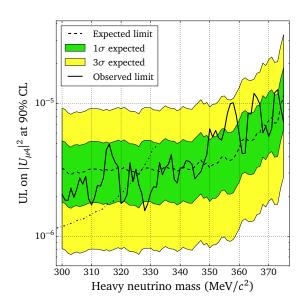


Figure 6: Expected and observed upper limits (at 90% CL) on the matrix element squared $|U_{\mu4}|^2$ at each assumed ν_h mass. The existing limit from KEK E089 [6] is also shown (dotted line). Below 300 MeV/ c^2 there is a limit of $\mathcal{O}(10^{-8})$ from BNL E949 [7], not shown.

8 Conclusions

A peak search has been performed in the missing mass spectrum observed in $K^+ \to \mu^+ N$ decays using part of the NA62 2007 dataset. Limits in the range 2×10^{-6} to 10^{-5} have been set on the mixing matrix element squared between muon and heavy neutrino states for assumed neutrino masses in the range 300–375 MeV/ c^2 . The result extends the range of masses for which upper limits have been set on the value of $|U_{\mu4}|^2$ by previous ν_h production experiments. Thanks to the design and excellent performance of the current NA62 setup [14], a substantial improvement in sensitivity is expected.

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References

- [1] R. N. Mohapatra and P. B. Pal, World Sci. Lect. Notes Phys. 72 (2004) 1.
- [2] R. N. Mohapatra et al., Rept. Prog. Phys. 70 (2007) 1757.
- [3] T. Asaka, S. Blanchet, and M. Shaposhnikov, Phys. Lett. B 63 (2005) 151.
- [4] T. Asaka and M. Shaposhnikov, Phys. Lett. B 620 (2005) 17.
- [5] R. Shrock, Phys. Lett. B **96** (1980) 159.
- [6] R. S. Hayano et al. [KEK E089 Collaboration], Phys. Rev. Lett. 49 (1982) 1305.
- [7] A. V. Artamonov et al. [BNL E949 Collaboration], Phys. Rev. D 91 (2015) 052001.
- [8] D. Gorbunov and M. Shaposhnikov, JHEP 0710 (2007) 015, Erratum JHEP 1311 (2013) 101.
- [9] J. R. Batley et al. [NA48/2 Collaboration], Eur. Phys. J. C 52 (2007) 875.
- [10] V. Fanti et al. [NA48 Collaboration], Nucl. Instrum. Meth. A 574 (2007) 433.
- [11] C. Lazzeroni et al. [NA62 Collaboration], Phys. Lett. B **719** (2013) 326.
- [12] C. Lazzeroni et al. [NA62 Collaboration], Phys. Lett. B 698 (2011) 105.
- [13] W. A. Rolke and A. M. Lopez, Nucl. Instrum. Meth. A 458 (2001) 745.
- [14] E. Cortina Gil et al. [NA62 Collaboration], 2017 JINST 12 P05025.